

# The Disc DIRC Čerenkov Detektor at PANDA

Benno Kröck   Avetik Hayrapetyan   Irina Brodski  
Klaus Föhl   Marko Zühlsdorf   Michael Düren  
Michael Sporleder   Oliver Merle   Peter Koch

Justus Liebig Universität Gießen, Germany

HANUC Lecture Week 2009, Torino

## Outline

### 1 Motivation

- FAIR
- PANDA

### 2 Disc DIRC

- Particle Identification
- Focussing Light Guides Design
- Time of Propagation Design (ToP)

### 3 Further Investigations

- Photon sensors
- Dispersion

FAIR

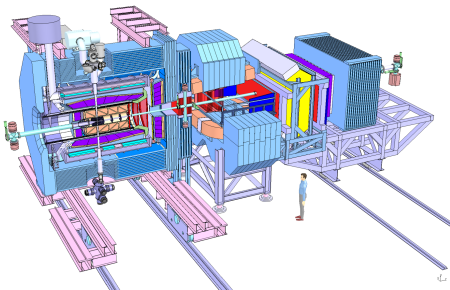
## Facility for antiproton and ion research



### New accelerator facility

- Gesellschaft für Schwerionenforschung (GSI)
- Darmstadt, Germany

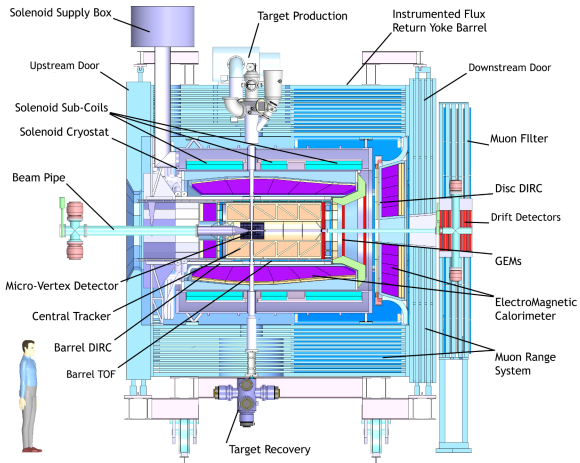
# Antiproton **an**ihilation at **D**armstadt



## Research

- Strong interaction, confinement
  - Structure of hadronic mass
  - Search for new hadrons: hybrid/glueball states
- 
- Fixed target
  - Antiproton beam of 1.5 . . . 15 GeV

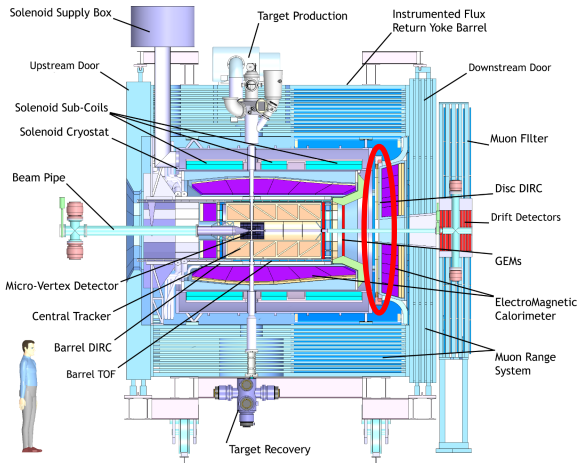
# Target spectrometer



► FSpec

- Little space for Čerenkov detector in forward region
- Disc DIRC in magnetic field

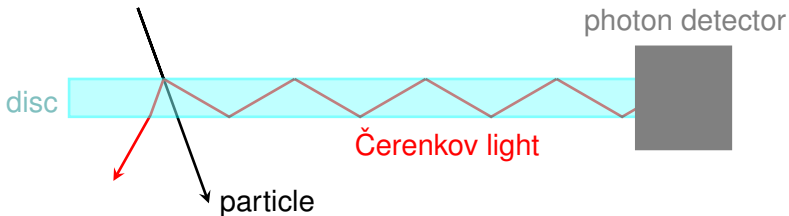
# Target spectrometer



► FSpec

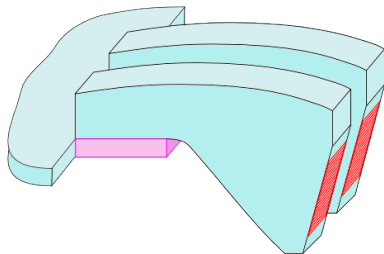
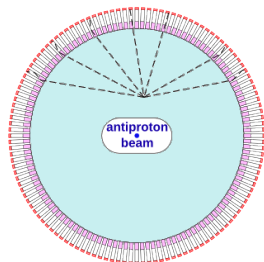
- Little space for Čerenkov detector in forward region
- **Disc DIRC** in magnetic field

## Disc detector of internally reflected Čerenkov light



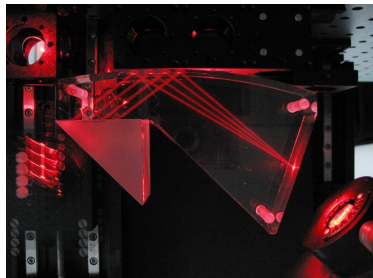
- Čerenkov light is emitted in disc (fused silica, 2 cm thick).
- Photons are propagated to detector by total internal reflexion.
- Particle identification possible, if momentum and velocity are known
  - Momentum from PANDA tracking detector
  - Velocity from Čerenkov angle  $\left(\cos \theta = \frac{1}{n\beta}\right)$

## Principle — focussing light guides



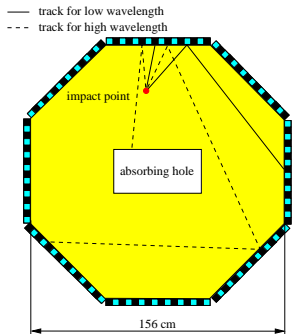
### Determination of Čerenkov angle

- First space coordinate from light guide position
- Second space coordinate from multi pixel photon detector





## Principle — time of propagation



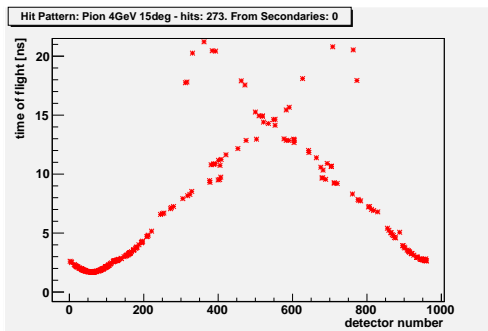
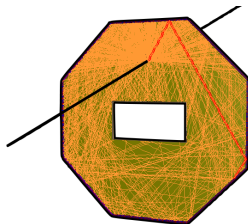
### Setup

- Octagonal disc
- Dichroic mirrors in alternating order between disc and sensors split wavelength spectrum into two ranges ( $\lambda < 500 \text{ nm}$  or  $\lambda > 500 \text{ nm}$ )
  - Reduction of dispersion effects
  - Increase of path lengths

### Determination of Čerenkov angle

- One space coordinate from position of photon detector
- Time of propagation

# Monte Carlo simulation

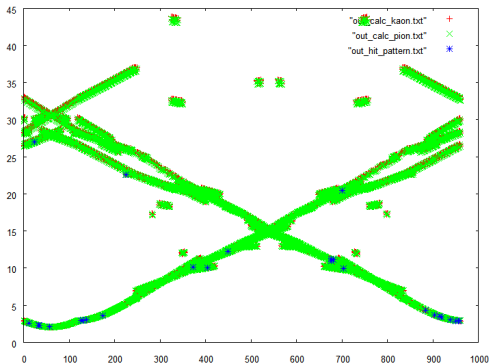


## 4 GeV pion

- $\approx 420$  Čerenkov photons emitted
- 273 photons propagated by total internal reflexion
- $\approx 70$  photons detected in reality (depending on photon detection efficiency)

Time of Propagation Design (ToP)

# Reconstruction

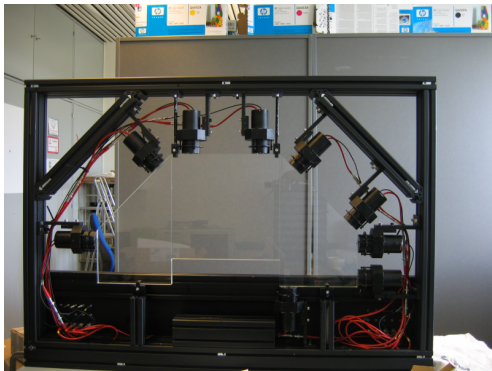


- No start signal
- Short and long paths are needed

- Calculation of time of propagation:
  - Kaon (red)
  - Pion (green)
- Comparison with simulation (blue) or finally measurement identifies particle.

Time of Propagation Design (ToP)

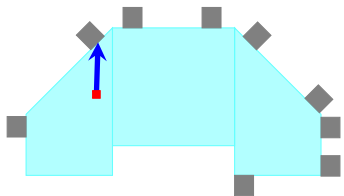
# Prototype



- We have a working prototype.
- Test beam experiment was made in October last year.
- Measurements with cosmics are running right now.

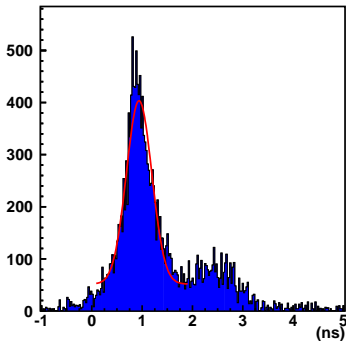
Time of Propagation Design (ToP)

## Results from cosmics measurements with prototype



Time of propagation  
calculated for relativistic  
particles:

- **0.92 ns**

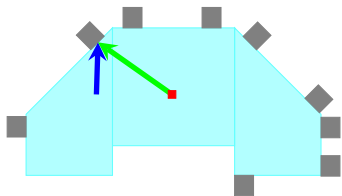


Measured:

- **0.94 ns**

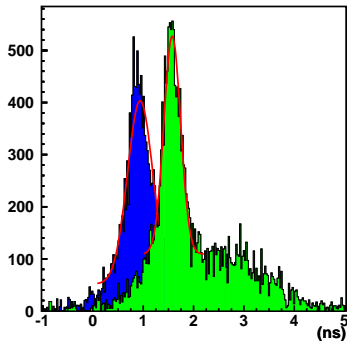
Time of Propagation Design (ToP)

## Results from cosmics measurements with prototype



Time of propagation  
calculated for relativistic  
particles:

- 0.92 ns
- 1.60 ns

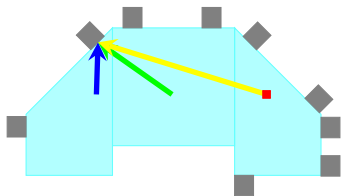


Measured:

- 0.94 ns
- 1.58 ns

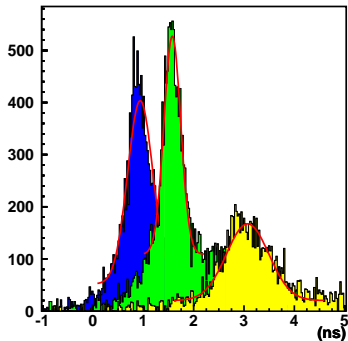
Time of Propagation Design (ToP)

## Results from cosmics measurements with prototype



Time of propagation  
calculated for relativistic  
particles:

- 0.92 ns
- 1.60 ns
- 3.12 ns

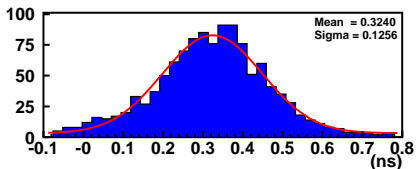
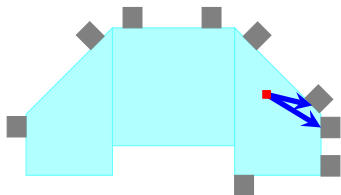


Measured:

- 0.94 ns
- 1.58 ns
- 3.07 ns

Time of Propagation Design (ToP)

## Results from cosmic measurements with prototype



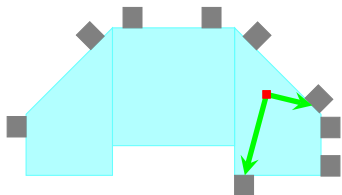
Calculated time differences:

- 0.40 ns



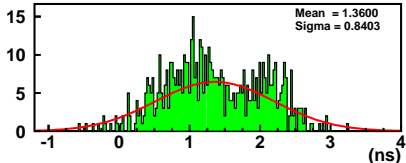
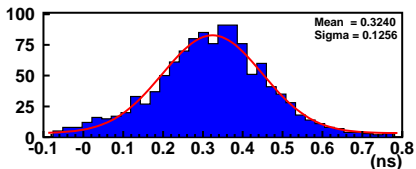
Time of Propagation Design (ToP)

# Results from cosmic measurements with prototype



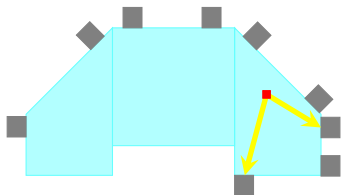
Calculated time differences:

- 0.40 ns
- 1.22 ns



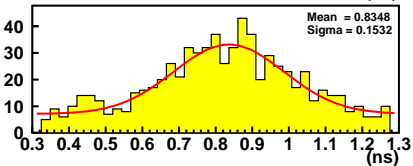
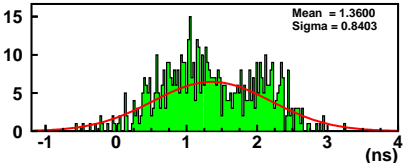
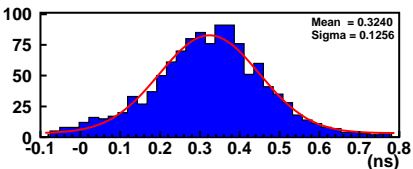
Time of Propagation Design (ToP)

# Results from cosmic measurements with prototype



Calculated time differences:

- 0.40 ns
- 1.22 ns
- 0.82 ns



## Requirements

- Very good time resolution ( $\sigma \approx 40$  ps)
- Low dark count rate
- Life time
- Operation in magnetic field ( $B \approx 2$  T)
- Radiation hardness

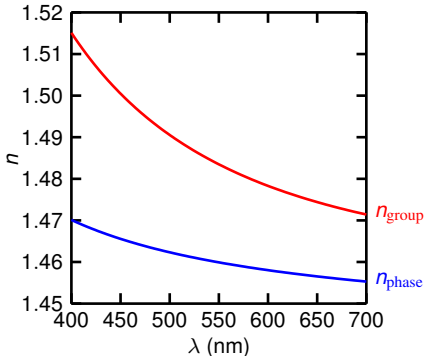
## Candidates

- Conventional PMTs: not working in magnetic field
- MCPMTs: short life time
- G-APDs: high dark count rate, eventually radiation damage

## Solutions

- MCPMTs with protection layer?
- Cooled G-APDs?
- Anything else?

## Effects in fused silica



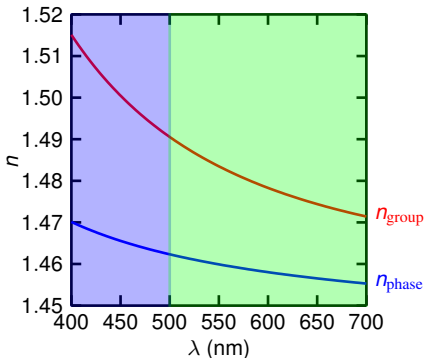
### At decreasing wavelength

- $n_{\text{phase}}$  increases. Čerenkov angle increases. Time of propagation decreases.
- $n_{\text{group}}$  increases. Group velocity decreases. Time of propagation increases.

### Further research to reduce dispersion effects

- Cutting out parts of wavelength spectrum
- Chromatic correction by achromatic units or software
- Splitting spectrum with dichroic mirrors as described above

## Effects in fused silica



### At decreasing wavelength

- $n_{\text{phase}}$  increases. Čerenkov angle increases. Time of propagation decreases.
- $n_{\text{group}}$  increases. Group velocity decreases. Time of propagation increases.

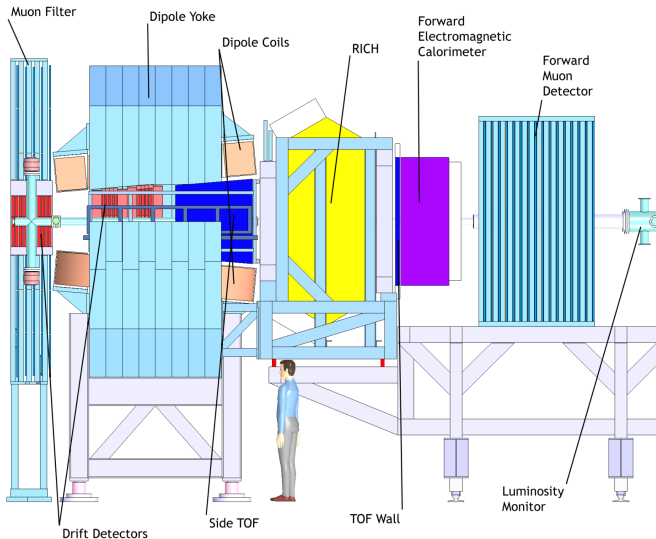
### Further research to reduce dispersion effects

- Cutting out parts of wavelength spectrum
- Chromatic correction by achromatic units or software
- Splitting spectrum with dichroic mirrors as described above

## Summary

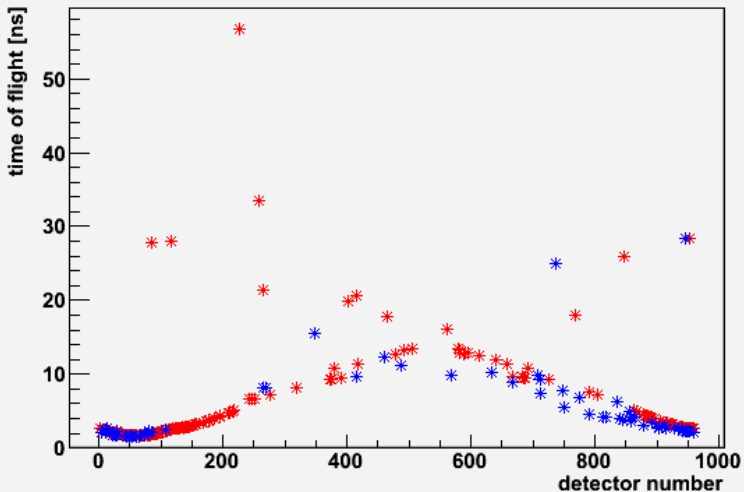
- Disc DIRC will provide particle identification at PANDA.
- Compact detector
- Two designs:
  - Focussing light guides
  - Time of Propagation

# Forward spectrometer



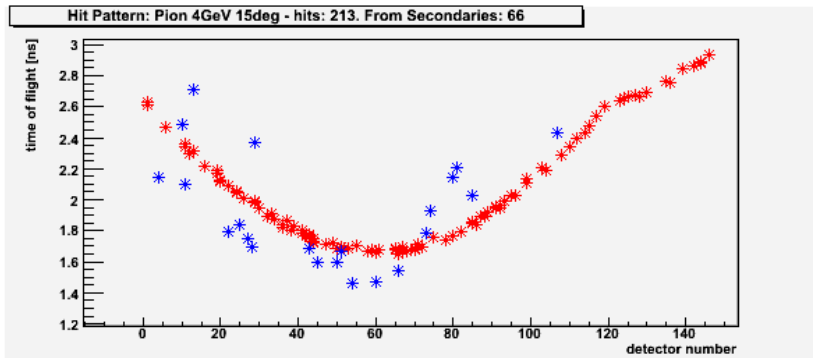
## Secondary events

Hit Pattern: Pion 4GeV 15deg - hits: 213. From Secondaries: 66

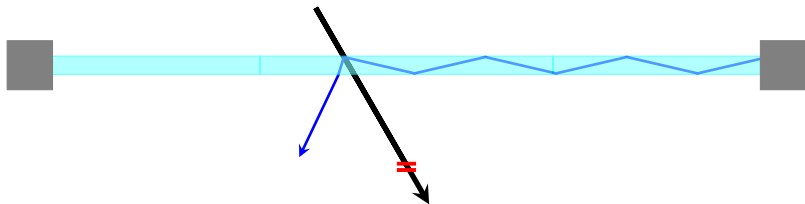




## Secondary events



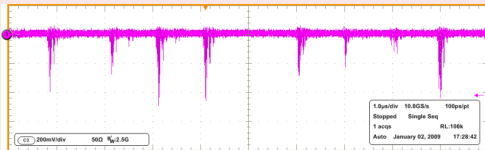
## Angle dependence



- Right sensor can see emitted cherenkov light.
- Left sensor does not see anything.

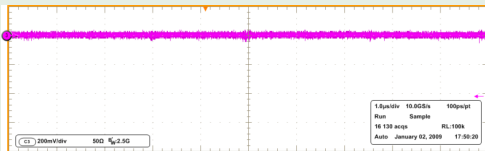
# Voltage supply

## Problem



Noise from high voltage modules

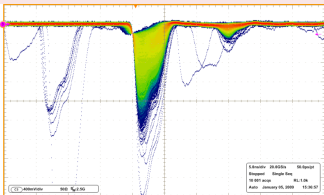
## Solution



Ferrites in feed lines

## Voltage divider boards

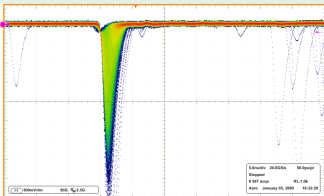
### Problem



Our boards were evil:

- Capacitances
- Long unshielded parts in signal cables

### Solution



- Voltage divider resistors directly mounted on the MCPs
- Signal cables directly connected to the MCPs